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Transverse Wave Induced Kelvin-Helmholtz rolls in Spicules ('twikhis')

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Outline

- Introduction
- IRIS & Hinode observations of spicules
 - Imaging & spectral properties
- Numerical modelling interpretation
 - Effects of a Transverse MHD wave on a spicule
- Discussion & Conclusions

Introduction



- Ubiquitous jets protruding from the chromosphere into the corona, type I-II
- I magneto-acoustic wave driven (ballistic, v<40 km/s) (Beckers 1968, Sterling 2000) II - Heating, fast upflow (v<110 km/s) (De Pontieu+ 2007, Rouppe vd Voort 2009, Sekse 2012, 2013).
- Multi-stranded, swaying and torsional motions (*Suematsu*+ 2008, *Pereira*+2012, *Skogsrud*+ 2014, *De Pontieu*+ 2014, *Rouppe v.d. Voort* 2015). Large transverse amplitudes (~10-20 km/s)
- Important transverse component. Which observational characteristics can be attributed to Alfvénic waves?

SST + IRIS

IRIS & Hinode Observations

- @North limb on 2014 April 29 01:00 02:00 UT (IHOP 249)
- IRIS sit-and-stare, 9.5 sec (SG), 19 sec (SJI 1400+2796)
- Hinode/SOT (Ca II H), 4.8 sec



IRIS/SJI 2796

- Transverse oscillation (1 period~200±20 sec)
- Coherent motion of strands (SOT) -> collective standing mode
- Disappearance in Ca II H (SOT)
- Monolithic structure (SJI 2796 & 1400)
- Doppler shift sign change at max displacement (Mg II k)
- Line width enhancement (Mg II k)

Case 1



- Case 2 ~ Case 3, but propagating disturbance
- Transverse oscillation
- Case 3: out-of-phase motion of strands (SOT)
- Disappearance in Ca II H (SOT)
- Monolithic structure (SJI 2796 & 1400)
- Case 3: Doppler shift sign change fixed along spicules axis (Mg II k)
- Line width enhancement (MgII k)

Cases 2 & 3



Cases 4 & 5

• Transverse oscillation



Numerical simulations

Numerical setup

- 3D MHD ideal simulations. CIP-MOCCT code (Kudoh et al. 1999).
 Adiabatic conditions.
- non-uniform grid, 1/4 tube (x,y,z)= (512, 256, 100) -> Δx,Δy=7.8 km, Δz=500 km
- Symmetric boundary conditions. Periodic in x

Forward modelling

- Optically thin approximation: FoMo code (Van Doorsselaere et al. 2016)
- Radiative transfer: 1.5-D RH code in PRD (Pereira & Uitenbroek 2015)

Loop & spicule model

- Straight loop, static spicule,
 L = 100 Mm, radius = 500 km,
 boundary layer width = 200 km
- Fundamental kink mode, P~ 255 s
- v₀ ~ 33 km/s (12 km/s at spicule top)





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Numerical results

S Stranded Si IV Intensity Mg II Intensity Mg II Dopp, vel.Mg II line width 45° LOS 45° LOS 45° LOS 45° LOS Cross-section at z=11 Mm structure in Voop Z[×10⁹cm⁻³] 50 40 40 density intensity, Doppler 10 0.5 <u>+9 km/s</u> < -9 km/s0.0 & line widths 5 • Apparent twist & 0.5 10.0 2×-0.5 -0.5 20 -0 [Log K] Temperature 6.0 -5 5.5 5.0 4.5 propagation 151 4.0 [km/s] [km/s] Height [Mm] 0.5 14 • Large amplitude $\frac{1}{2}$ 10 0.0 -0.5Doppler shifts Line 10 1 width 3 0.3 0.2 0.1 0.5 [Mm] -Vorticity Enhanced line 0.0 8 0.0 -0.1 -0.2 -0.3 -0.5 widths 5 в [cgs] Current density 0.5 ● Fast appearance /ॾ 0,5 0.0 0.0 disappearance of $\overline{>}$ -0.5 -0,5 strands -0.50 0.50 [km/s] × [Mm] perp. to LOS [Mm] perp, to LOS perp, to LOS perp, to LOS [Mm] [Mm] [Mm]

Numerical results



- Loop body visible in Mg II, but invisible in Ca II H & Si IV
- Bursty intensity profiles, particularly in Ca II H
- Doppler shift change at max displacement, but ragged: TWIKH rolls
- Maximum Doppler values at edges (out-of-phase with POS motion, *Okamoto*+2015, *Antolin*+2015)

Simulations

Si IV Intensity Mg II Intensity

RT Ca II Intensity

-0.5 0.0 0.5

perp. to LOS [Mm]

RT

Mg II Intensity

-0.5 0.0 0.5

perp. to LOS [Mm]

15

10

600

Time [s] 005

200

Height [Mm]

In agreement...

Observations

- Collective motion
- Doppler shift sign change at max displacement, with ragged transition
- Line width enhancement

RT Mg II Dop. vel.

-0.5 0.0 0.5

perp. to LOS [Mm]

- Loop body visible in Mg II k
- Fast appearance / disappearance of Ca II H strands

RT Mg II line width

-0.5 0.0 0.5

perp. to LOS [Mm]

-5

Line width

ˈkm/s

What is the spicule? The individual strand or the collection?



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velocity [km/s]

Doppler

_ine width [km/s]

14 12



- Collective motion suggests existence of larger flux tube
- Fast appearance / disappearance of strands in Doppler & line width maps
- Increase of amplitude with height -> apparent upward propagation

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Simulations

Observations

20

14

3

-2

19

17

14

12

Line width [km/s]

velocity [km/s]

Doppler

IRIS 1400



In disagreement...

- No strong temperature rise from KHI mixing (Magyar & Van Doorsselaere +2016, Karampelas+ 2017)
- Little intensity enhancement in Si IV
- No fast longitudinal speeds





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Time 1958 s

IRIS 2796

2014-04-29 UT01:33:09

SOT Ca II

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- Rotation: Doppler shift transition fixed in space & out-of-phase motion of substructure: match cases 3, 4 & 5 (similar to *Okamoto*+ 2016)
- Also propagating long wavelength torsional Alfvén wave? However, periodic asymmetry for half/part of flux tube is not observed. Also, sum over multiple shells of constant Alfvén speeds should decrease Doppler signal

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Conclusions

- IRIS&Hinode/SOT Observation of spicules: 2 major cases
 - Little coherence among strands and Doppler shift sign change along spicule axis -> *rotation or long wavelength torsional Alfvén waves*?
 - Coherent motion of strand sub-structure & Doppler shift sign change at maximum displacement -> kink waves?
- 3-D MHD model of spicule/loop oscillating with a transverse MHD wave -> TWIKH rolls:
 - Strand-like structure in intensity, Doppler shifts and line widths
 - Coherent motion of sub-structure on average
 - Doppler shift sign change (ragged) at maximum displacement & increased line widths
 - Strong variations in Doppler shifts, line widths & Ca II H intensity due to KHI mixing: agreement with observations
 - Small increase in Si IV: disagreement. -> existence of additional mechanism at work in spicules

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